# Image Formation and Capture

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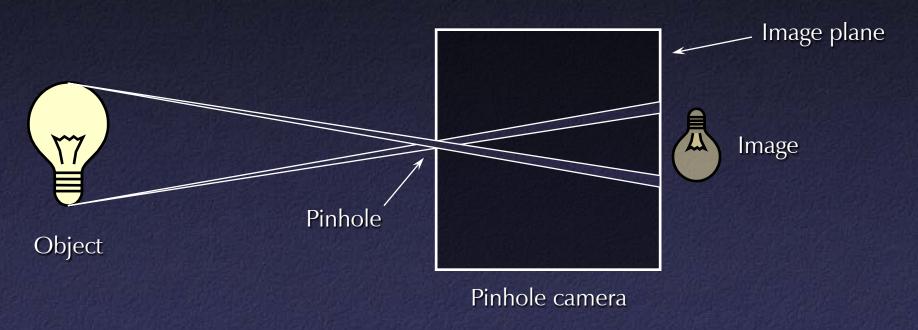
- Devices
- Sources of Error

## Optics

- Pinhole camera
- Lenses
- Focus, aperture, distortion

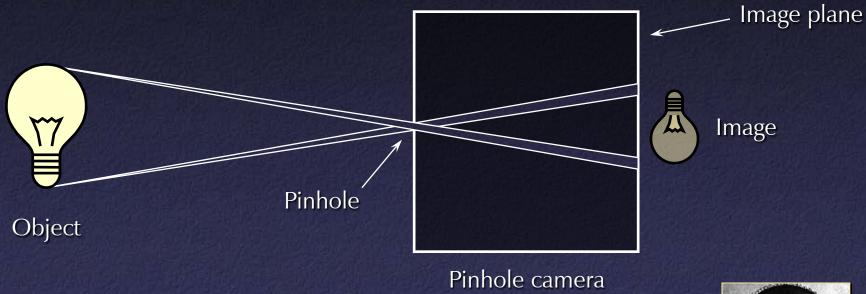
## Pinhole Camera

"Camera obscura" – known since antiquity

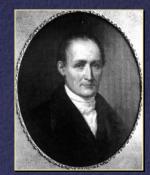


#### Pinhole Camera

"Camera obscura" – known since antiquity



 First recording in 1826 onto a pewter plate (by Joseph Nicéphore Niepce)

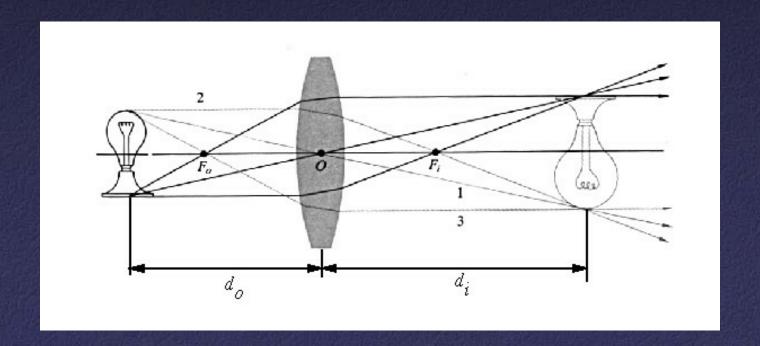


#### Pinhole Camera Limitations

- Aperture too big: blurry image
- Aperture too small: requires long exposure or high intensity
- Aperture much too small: diffraction through pinhole ⇒ blurry image
  - Rule of thumb: aperture should be significantly larger than wavelength of light (400-700 nm)

#### Lenses

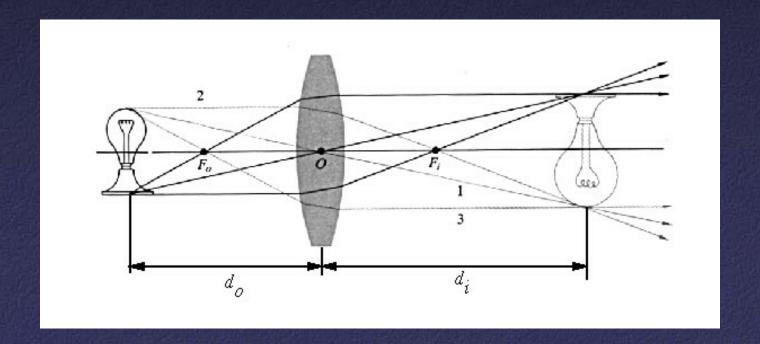
- Focus a bundle of rays from a scene point onto a single point on the imager
- Result: can make aperture bigger



## Ideal "Thin" Lens Law

 Relationship between focal distance and focal length of lens:

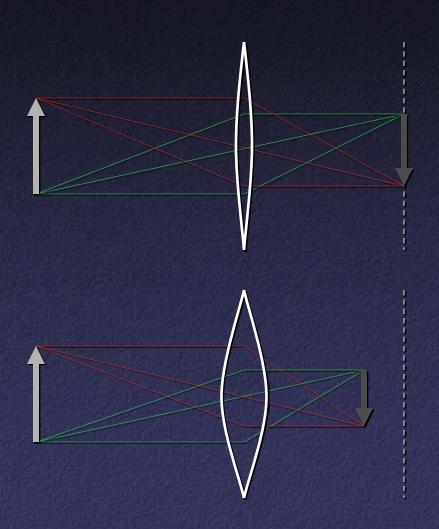
$$1/d_o + 1/d_i = 1/f$$



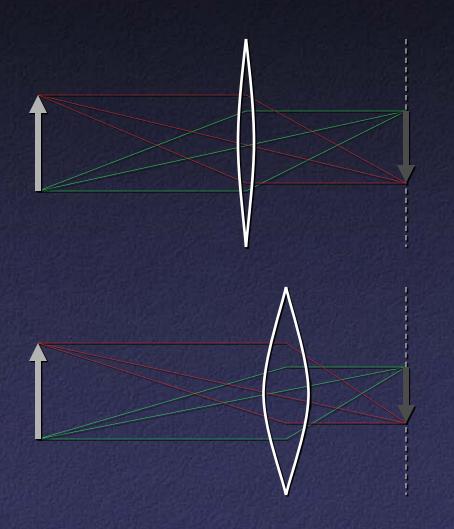
# Camera Adjustments

- Focus?
  - Changes  $d_i$
- Iris?
  - Changes aperture
- Zoom?
  - Changes f and sometimes  $d_i$

# Zoom Lenses – Varifocal

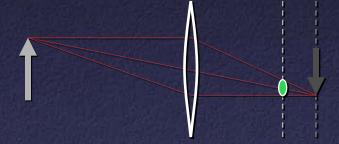


# Zoom Lenses – Parfocal



# Focus and Depth of Field

- For a given  $d_i$ , "perfect" focus at only one  $d_o$
- In practice, OK for some range of depths
  - Circle of confusion smaller than a pixel

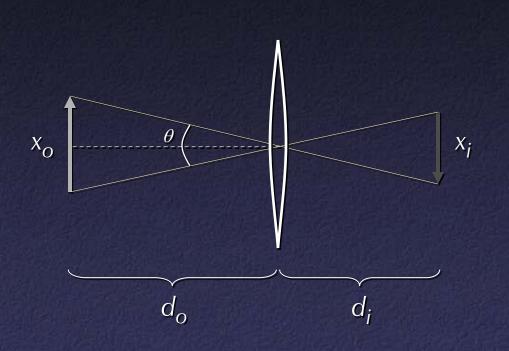


- Better depth of field with smaller apertures
  - Better approximation to pinhole camera
- Also better depth of field with wide-angle lenses

## Field of View

- Q: What does field of view of camera depend on?
  - Focal length of lens
  - Size of imager
  - Object distance?

# Computing Field of View



$$1/d_o + 1/d_i = 1/f$$

$$\tan \theta/2 = \frac{1}{2} x_o/d_o$$

$$x_o/d_o = x_i/d_i$$

$$\theta = 2 \tan^{-1} \frac{1}{2} x_i (1/f - 1/d_o)$$

Since typically 
$$d_o >> f$$
,  
 $\theta \approx 2 \tan^{-1} \frac{1}{2} x_i / f$   
 $\theta \approx x_i / f$ 

## Aperture

- Controls amount of light
- Affects depth of field
- Affects distortion (since thin-lens approximation is better near center of lens)

## Aperture

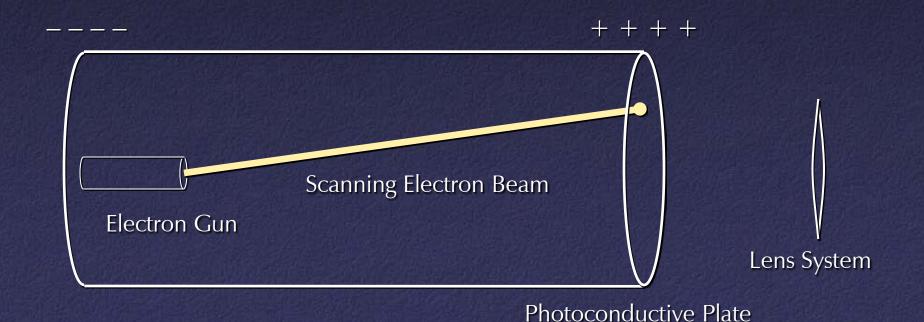
- Aperture typically given as "f-number"
- What is *f*/4?
  - Aperture diameter is 1/4 the focal length
- One "f-stop" equals change of f-number by  $\sqrt{2}$ 
  - Equals change in aperture area by factor of 2
  - Equals change in amount of light by factor of 2
  - Example:  $f/2 \rightarrow f/2.8 \rightarrow f/4$

## Sensors

- Film
- Vidicon
- CCD
- CMOS

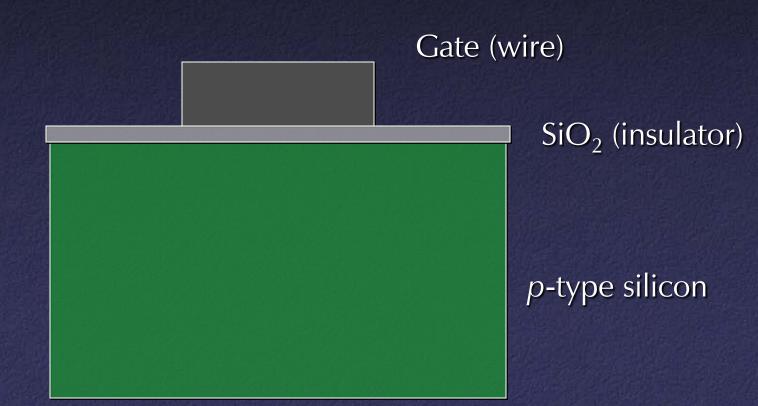
## Vidicon

- Best-known in family of "photoconductive video cameras"
- Basically television in reverse



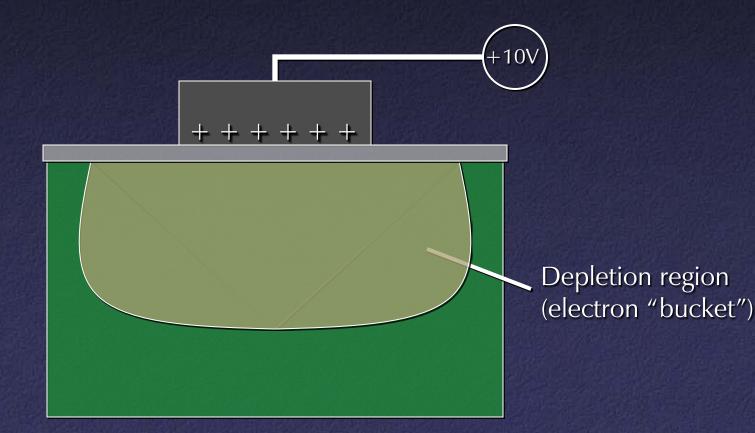
# MOS Capacitors

MOS = Metal Oxide Semiconductor



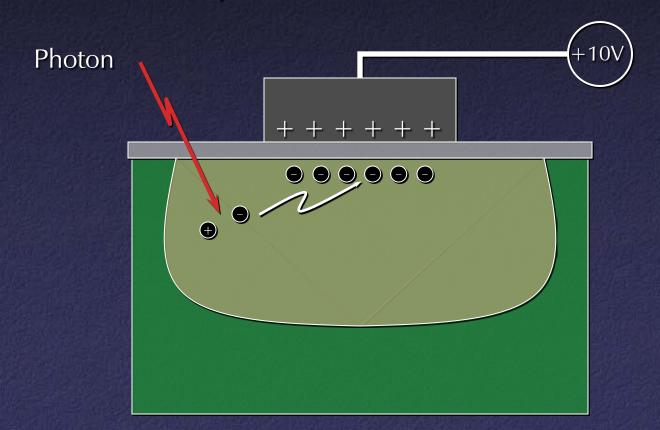
# MOS Capacitors

 Voltage applied to gate repels positive "holes" in the semiconductor



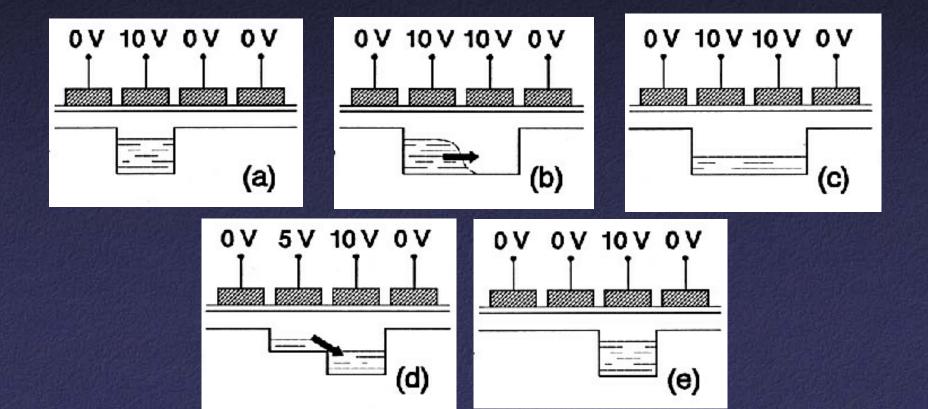
# MOS Capacitors

 Photon striking the material creates electron-hole pair



## Charge Transfer

 Can move charge from one bucket to another by manipulating voltages

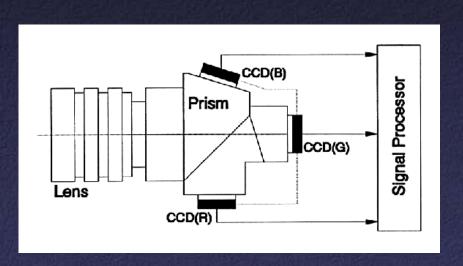


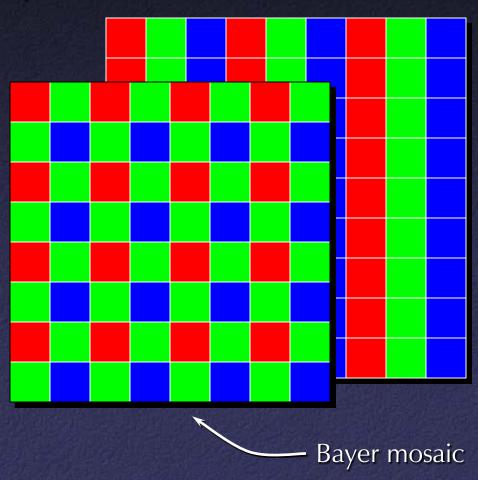
## CMOS Imagers

- Recently, can manufacture chips that combine photosensitive elements and processing elements
- Benefits:
  - Partial readout
  - Signal processing
  - Eliminate some supporting chips  $\Rightarrow$  low cost

## Color

• 3-chip vs. 1-chip: quality vs. cost





# Errors in Digital Images

• What are some sources of error in this image?

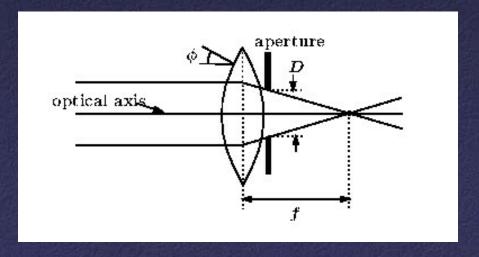


## Sources of Error

- Geometric (focus, distortion)
- Color (1-chip artifacts, chromatic aberration)
- Radiometric (cosine falloff, vignetting)
- Bright areas (flare, bloom, clamping)
- Signal processing (gamma, compression)
- Noise

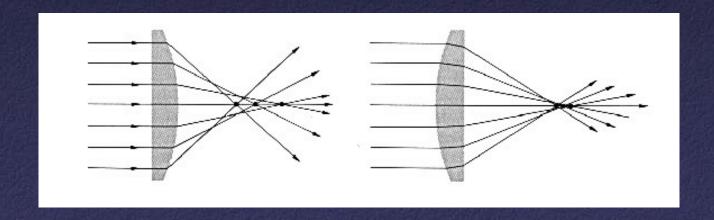
## Monochromatic Aberrations

- Real lenses do not follow thin lens approximation because surfaces are spherical (manufacturing constraints)
- Result: thin-lens approximation only valid iff  $\sin \varphi \approx \varphi$

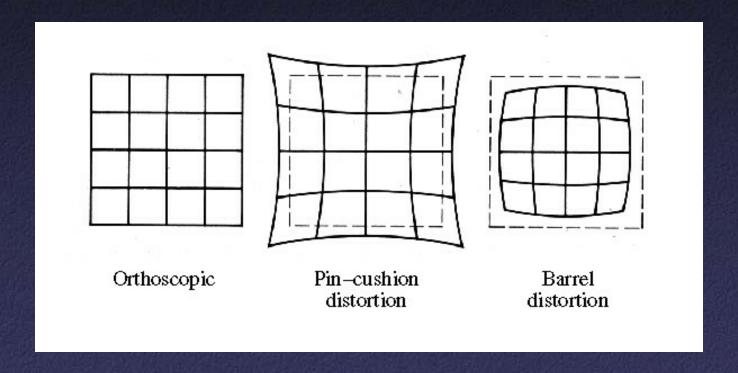


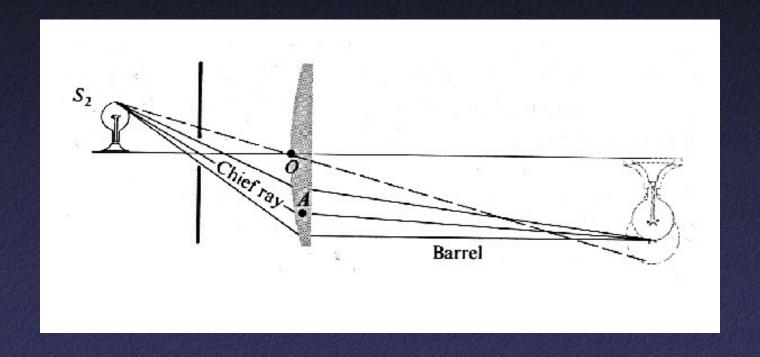
# Spherical Aberration

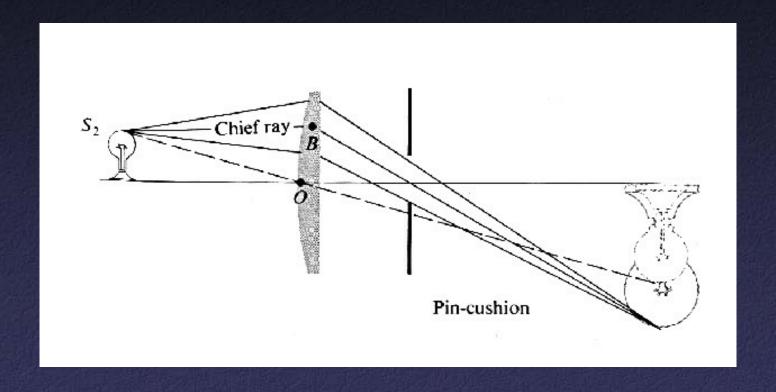
- Results in blurring of image, focus shifts when aperture is stopped down
- Can vary with the way lenses are oriented

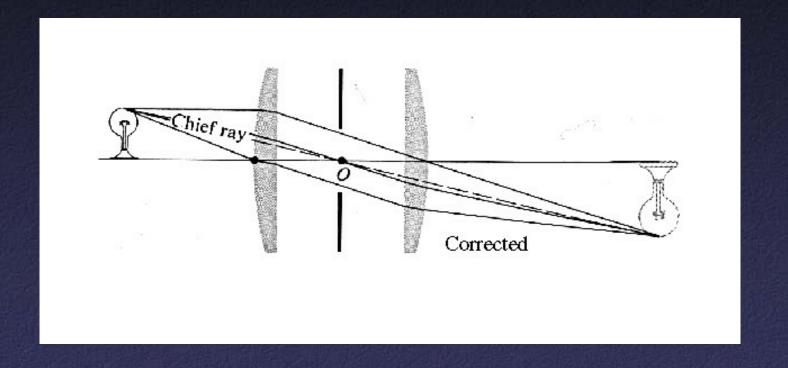


Pincushion or barrel radial distortion









## First-Order Radial Distortion

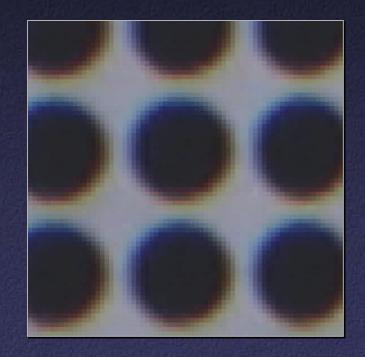
- Goal: mathematical formula for distortion
- If small, can be approximated by "first-order" formula (like Taylor series expansion):

$$r' = r (1 + \kappa r^2)$$
  
 $r = \text{ideal distance to center of image}$   
 $r' = \text{distorted distance to center of image}$ 

Higher-order models are possible

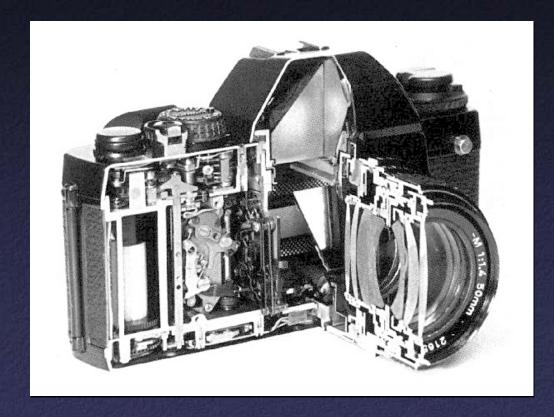
## Chromatic Aberration

- Due to dispersion in glass (focal length varies with the wavelength of light)
- Result: color fringes
- Worst at edges of image
- Correct by building lens systems with multiple kinds of glass



## Correcting for Aberrations

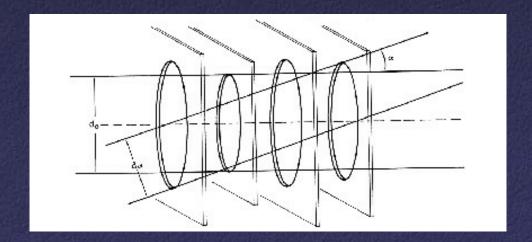
High-quality
 compound lenses
 use multiple
 lens elements to
 "cancel out"
 distortion and
 aberration



Often 5-10 elements, more for extreme wide angle

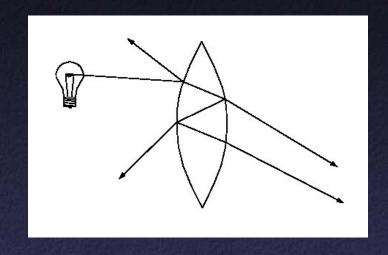
## Other Limitations of Lenses

- Optical vignetting: less power per unit area transferred for light at an oblique angle
  - Transferred power falls off as  $\cos^4 \varphi$
  - Result: darkening of edges of image
- Mechanical vignetting: due to apertures



## Other Limitations of Lenses

Flare: light reflecting
 (often multiple times)
 from glass-air interface



- Results in ghost images or haziness
- Worse in multi-lens systems
- Ameliorated by optical coatings (thin-film interference)

#### Bloom

- Overflow of charge in CCD buckets
  - Spills to adjacent buckets
  - Streaks (usually vertical) next to bright areas
- Some cameras have "anti-bloom" circuitry

# Flare and Bloom



Tanaka

# Dynamic Range

- Most common cameras have 8-bit (per color channel) dynamic range
  - Can be nonlinear: more than 255:1 intensity range
- Too bright: clamp to maximum
- Too dim: clamp to 0
- Specialty cameras with higher dynamic range (usually 10-, 12-, and 16-bit)

# High Dynamic Range (HDR) from Ordinary Cameras

- Take pictures of same scene with different shutter speeds
- Identify regions clamped to 0 or 255
- Average other pixels, scaled by 1 / shutter speed
- Can extend dynamic range, but limitations of optics and imager (noise, flare, bloom) still apply

#### Gamma

- Vidicon tube naturally has signal that varies with light intensity according to a power law: Signal =  $E^{\gamma}$ ,  $\gamma \approx 1/2.5$
- CRT (televisions) naturally obey a power law with gamma ≈ 2.3–2.5
- Result: video signal standard has gamma of 1/2.5
- CCDs and CMOS linear, but gamma ≈ 2.2 almost always applied

## Noise

- Thermal noise: in all electronics
  - Noise at all frequencies
  - Proportional to temperature
  - Special cooled cameras available for low noise
- Shot noise: discrete photons / electrons
  - Shows up at extremely low intensities
  - CCDs / CMOS can have high efficiency approaching 1 electron per photon

## Noise

- 1/f noise inversely proportional to frequency
  - Not completely understood shows up in semiconductors
  - Can be dominant source of noise
- All of the above apply for imager and amplifier

# Filtering Noise

- Most common method simple blur
  - e.g., convolution with Gaussian
- Adaptive filters to prevent bleed across intensity edges
- Other filters for specialized situations
  - e.g., "despeckling" (median filters) for dead pixels